The EIS Planning Tool User's Guide

The EIS Planning Tool User's Guide Introduction About This Guide Software Requirements Structure of a Study Line Lists **Rasters Studies** A Quick Word on the Timeline **Preparing a Raster** Starting the Raster Tool Making a Line List Adding a Project Line Adding a Custom Line Removing a Line from the Line List Saving a Line List Exporting a Line List Importing a Line List Changing the CCD image Making a Raster Changing the type of raster Making a Study Notes

Introduction

About This Guide

A planning guide will develop here over the coming weeks. As our understanding of the instrument grows, this guide will inevitably — and desirably — develop. This naturally lends itself to a TWiki platform, although a PDF version is planned for each significantly new version.

N.B. This is a work in progress. Please ignore until it is linked from the SolarB WebHome

Software Requirements

In order to run the EIS planning software, you will need:

- 1. IDL; the planning tool is developed under Linux, but works under Windows XP and Mac OS X, too.
- 2. an up-to-date installation of *solarSoft*, including the EIS branch of the SSW tree.

Structure of a Study

An EIS *study* is composed of one or more *rasters*, each of which is defined in part by a *line list*. To those with past experience of <u>preparing a SoHO CDS study</u>, this nomenclature may well be familiar. However, it is worth explaining both for those who are new to the concept of studies, **and** for those new to EIS.

Line Lists

A line list is simply a list of wavelengths (in *angstroms*) at which the instrument is to observe, along with an identifier (or name) for that wavelength (typically, the ion species which produces it, but this can be customised). No information about the spectral range around

each wavelength is assumed.

For example, the core line list of EIS could be expressed as follows:

Name	Wavelength
Ca XVI	I 192.82
Fe XII	195.12
He II	256.32

This line list thus consists of three lines, with ion names for each.

Rasters

In EIS nomenclature, the term raster covers both possible forms of observation:

- Scanning rasters, which are made by building up slit (or slot) images across an area from west to east, thereby achieving spatial coverage in the west-east direction (the slit of EIS is oriented the north-south direction). Such observations can be thought of as having dimensions of *x*, *y* and λ.
- Sit & stare observations, which are successive slit/slot exposures at the same nominal position on the Sun. These are observations in *x*, *y* and *t*.

Altogether, the definition of a raster is determined by several parameters:

- 1. which line list to use
- 2. which slit/slot to observe through
- 3. the width of each spectral window
- 4. what exposure duration(s) should be used
 - a. more than one exposure time can be specified: this is useful if, for example, more than one lines is to be observed, but one is much stronger than the rest.
- 5. whether it is a scanning or sit'n'stare raster
 - a. if it is a scanning raster:
 - i. how many steps to make in the x direction
 - ii. how wide these steps should be
 - b. if it is a sit'n'stare raster
 - i. how many times the exposure(s) should be repeated
- 6. a unique name, or "acronym", such as NRL_FLARE_EVAP.

Studies

A study* can be most simply defined as consisting of one or more rasters grouped together in a particular order for some science objective.

For a study to be completely defined, we must specify

- 1. the order in which component rasters are to be run
- 2. how often each raster is to be repeated
- 3. what instrument triggers will be enabled/acknowledged
- 4. what compression scheme (if any) is to be used
- 5. a unique acronym, as is also true for rasters (see above)

(*Note that, to reduce confusion of terminology between "study" and "sequence" within EIS, the term "sequence" refers only to low-level engineering commands.)

A Quick Word on the Timeline

Again, this concept will be familiar to former CDS planners, but it is relatively self-explanatory to others, too.

The **timeline** is the interface which allows the scheduling of studies for a given 24-hour period. However, as this will currently only be used by EIS Chief Observers (COs), we won't go into it in this guide as yet.

Preparing a Raster

Starting the Raster Tool

Once you have started your *solarsoft IDL* session, you can begin creating your raster by typing *eis_mk_raster*. You should now be presented with a graphical user interface (or "widget" in *IDL* speak) that looks something like this:



This represents a default raster (<u>Raster ID #000001</u>), which is loaded in when the raster tool is started. You'll see that it contains just the three core lines in its line list (7). On the left is the Line List area (2):

- (3) the Linelist ID (<u>#000001</u>)
- (4) the Linelist name, or **acronym** (Core_Lines)
- (5) the date the LineList was saved, and
- (6) the person who saved it. In this case, Matt.

The panel to its right (7) contains the most scientifically important information, namely:

- the wavelength of interest (in effect, this is the central wavelength of the spectral range or "window" that you're interested in)
- the name of that window (typically the ion of interest)
- the CCD on which each wavelength falls
 - (16) CCD 1: 170 210 Å
 - (17) CCD 2: 230 270 Å
- the width (in pixels) of each spectral window

The information on the far right, under **Selected Line** (8) is initially greyed-out because it is line-specific, and no line has yet been selected. Click on the first line in the list (Ca XVII), and you should see this area activated. You should also see the corresponding spectral window highlighted in yellow in the synthetic CCD 1 image (16) below.

	_	_	_	E D	and the state	ster to		-			
Exit Linelist Raster Plot Edit								8			
Line List ID: 000001 Acronym: Core_Lines Created: 26-Apr-2006 Author: Matthew Whillock	# 2 3	Wavelengt 192.82 195.12 256.32	th(Å) (c) (c) (c)	Name Ca XVII Pe XII He II	Pixel 1189 1292 472	CCD 1 2	Width 32 32 32 11 12	Selected Pixels: Wavelength: Width: Name:	Line 1173 192.47 32 * Spect • Text (: [1189 : [192.82 tral notation (custom)	: 1204 -
 Quiet > Active > Flare Cold 100 200 300 400 170 172 174 17 	our Table:	600	20-PUR	8PLE 800 90	- 100 100	0 11	00 1200	1300 1400 94 196 196	1500	1600 17/ 202 204	13 20 1800 206 20
		5 100	102	104 100			192				
100 200 300 400 Silt/slot selection 1 Window height (pixels) 512	500	600 Raster Type	700 700 Defa	800 90 * Scanning ault width (pixe	o too Sit'n	'Stare	192	1300 1400 Set all	1500	1600 170	00 1800
100 200 300 400 Slit/slot selection 1 Window height (pixels) 512 Number of exposures/step 1 Number of Fine Mirror steps 1/40	soo Step	600 Raster Type t Exposure/D size	700 a: Defa Oelay Ti	eoo eo Scanning ault width (pixe imes Exp lumber of expo	e too Sit'n els) osure(Del	*Stare 32 ====================================	192 100 1200 e s(ms) 50.	1300 1400 Set all 00(10)	1500	1600 170	00 1800
100 200 300 400 Silt/slot selection 1 Window height (pixels) 512 Number of exposures/step 1 Number of Fine Mirror steps [40] Raster ID: 000001 Acroo Duration: 35m33s Volu Title: [RA:000001] [Scare	500 500 Step Step mme: 314 n:40(2*)	600 Raster Type t Exposure/D size [2 fault Rast 488 kBits steps][ss:	700 Befay Ti (N :er	eoo eo Scanning ault width (pixe imes Exp lumber of expo Author Data ra H: 512, nWins	o tioc Sit'n els) osure(Del isures = 4 Jot isures = 15 :3, LL : 00	•Stare 32 [] ay) time 1) 1113 1 0001][e s(ms) 50. Bits/s ExpT(Delay	1300 1400 Set all 00(10) Created: 26-30	1500 1500	1600 17	oo 1800 Solar-F EIS

The **Selected Line** section will now let you set a number of parameters related to the spectral window. The area labelled **Pixels** (9) lets you adjust the position of the centre of the window on the detector. Alternatively you can do this in wavelength (10). In either case, the numbers to the left and right of the field represent the lower and upper limits to the spectral window (in CCD pixel and wavelength units, respectively). Before launch, the translation between pixel co-ordinates and angstroms is done using a simple linear relation. After EIS is commissioned, this will be calibrated using real data, and the wavelength calibration will be regularly monitored and updated.

Also, note that — as with all editable IDL text widget fields, **YOU MUST PRESS ENTER ONCE YOU HAVE TYPED THE VALUE YOU WANT**; otherwise, the widget will simply not process your desired value and your efforts will go unnoticed by the system.

To adjust the spectral width of the window, use the slider (11) to choose a value, measured in units of pixels. For on-board software reasons, you can only choose integer multiples of 8 pixels. The default value is 32 pixels, which is pretty generous: in the middle of CCD 1 (190 Å) this is about \pm 400 km s⁻¹; in the middle of CCD 2, around \pm 250 km s⁻¹. But you may want to increase or decrease this for science or telemetry reasons.

Another option, although you probably won't often want to do this, is to select half a CCD as your spectral window. This is the biggest possible window, since EIS has two read-out nodes per CCD, located at opposite ends of the CCD: it therefore normally reads only half a CCD from each node.

_If you want to adjust all the spectral windows at once to the same width, read below... _

Finally you have the option of changing the name of the spectral window. If you're dealing with a core line, then you don't need to change this. However if you've added a line, particularly a custom line, you should check that it is appropriately named.

Making a Line List

The first step in creating an observation is to decide in which wavelengths — or which wavelength ranges — you want to observe.

Be aware that the core line list <u>above</u> **must** be included in all observations. This is because standard data products (line-intensity, doppler-shift, and line-width maps — known as Level 2 FITS) will be made from all EIS observations in a data pipeline; these products will be used to browse data in the online catalogue. (If you don't include the core lines in your linelist, you'll have it sent back to you by the **Scientific Schedule Co-ordinators (SSCs)**.) These core lines are marked **(c)** in (7).

Adding a Project Line

The EIS mission has certain expected lines from which you can choose to help you put together a line list. To add one of these **project lines** to your line list (which is probably the first thing you'll want to try), go to the main menu (1) and choose *Edit, Add a line from the EIS project list...*

The main linelist menu



This opens up a new window with a list of wavelengths and the ions which emit them:

By default, these known lines are ordered by wavelength, but you can change that by clicking on **Ascending Wavelength** and changing it to, for example, **Ascending Name**. This orders the ion species alphabetically (not by atomic number or roman numeral order, but its' enough to let you search for the ion you want):

P P	Atomic Name Fe XI S X	- 12
p p	Fe XI S X	14
р	SX	
14		
P	Fe X	
р	Fe VIII	
P	Ni XVI	
р	Fe VIII	
р	Fe XII	
р	Fe XII	
p	Fe XXI	
р	Ar XIV	
P	Fe XI	
p	Fe X	
р	Fe XII	
р	Ar XIV	1.61
р	Fe XXIV	
р	Fe XII	
с	Ca XVII	
р	0 V	
р	Fe XII	
р	Ca XIV	
P	Ni XVI	
p	Ar XIV	
р	Fe VIII	
с	Fe XII	
р	Ni XVI	
P	Fe VIII	
p	Fe XIII	
p	Fe XII	
p	SX	
	Pe VIII	1.00
	, , , , , , , , , , , , , , , , , , ,	p Fe VIII p Fe VIII p Fe VIII p Fe VIII p Fe XII p Fe XII p Fe XI p Fe XI p Fe XII p Ni XVI p Fe VIII c Fe XII p Fe VIII p Fe XIII p S X

EIS Line Selection Tool 🛛 🔍						
Wavelength	Type	Atomic Nam	e			
187.970	р	Ar XIV	11			
191.400	p	Ar XIV				
194.390	p	Ar XIV				
192.820	c	Ca XVII				
193.870	р	Ca XIV				
200.980	p	Ca XV				
180.400	p	Fe XI				
184.540	p	Fe X				
185.210	p	Fe VIII				
186.600	Р	Fe VIII				
186.850	р	Fe XII				
186.880	P	Fe XII				
187.890	р	Fe XXI				
188.230	р	Fe XI	1.1			
190.040	р	Fe X				
191.050	р	Fe XII				
192.040	р	Fe XXIV				
192.390	р	Fe XII				
193.520	р	Fe XII				
194.660	р	Fe VIII				
195.120	с	Fe XII				
195.970	p	Fe VIII				
196.540	р	Fe XIII				
196.650	р	Fe XII				
200.020	р	Fe XIII				
201.130	р	Fe XIII				
202.040	р	Fe XIII				
203.830	р	Fe XIII				
204.650	р	Fe XVII				
247.160	р	Fe XXII	57			
Sort: As	cending	j name 🖂]			
Add	Dismiss					

You're not limited to selecting one line at a time, either.

To select multiple lines, click and drag over a range

To select non-adjacent lines, hold down the **CTRL** key, and left-click on

But we'll just concentrate on adding a single line of ca xv for now (you can also add a single line by

double-clicking on it).

1	Vavelength 200.980 180.400 184.540 185.210 186.600 186.850 186.850 186.880 187.890 188.230 190.040 191.050 192.040	Type P P P P P P P P P P P P P	Atomic Name Ca XV Fe XI Fe X Fe VIII Fe VIII Fe XII Fe XII Fe XI Fe X Fe X Fe X		Vavelengi 187.970 191.400 194.390 192.820 193.870 200.980 180.400 184.540 185.210	h Type P P c P P P P P P	Atomic Name Ar XIV Ar XIV Ar XIV Ca XVII Ca XIV Ca XV Fe XI Fe X Fe X Fe VIII
	200,980 180,400 184,540 185,210 186,800 186,880 187,890 187,890 187,890 187,890 190,040 191,050 192,040	P P P P P P P P P	Ca XV Fe XI Fe X Fe VIII Fe VIII Fe XII Fe XII Fe XI Fe X Fe X Fe X		187.970 191.400 194.390 192.820 193.870 200.980 180.400 184.540 185.210	p p c p p p p	Ar XIV Ar XIV Ar XIV Ca XVII Ca XIV Ca XV Fe XI Fe X Fe X Fe VIII
	180,400 184,540 185,210 186,600 186,850 186,880 187,330 187,330 188,230 189,040 190,040 191,050 192,040	р Р Р Р Р Р Р	Fe XI Fe X Fe VIII Fe VIII Fe XII Fe XII Fe XI Fe XI Fe X Fe X		191.400 194.390 192.820 193.870 200.980 180.400 184.540 185.210	p p c p p p p	Ar XIV Ar XIV Ca XVII Ca XIV Ca XIV Fe XI Fe X Fe VIII
	184.540 185.210 186.600 186.850 186.880 187.890 188.230 190.040 191.050 192.040	P P P P P P P	Fe X Fe VIII Fe VIII Fe XII Fe XII Fe XI Fe X Fe X Fe X		194.390 192.820 193.870 200.980 180.400 184.540 185.210	P C P P P P	Ar XIV Ca XVII Ca XIV Ca XIV Fe XI Fe X Fe VIII
	185,210 186,600 186,850 186,880 187,890 188,230 190,040 191,050 192,040	р Р Р Р Р Р	Fe VIII Fe VIII Fe XII Fe XII Fe XI Fe X Fe X	0	192.820 193.870 200.980 180.400 184.540 185.210	c P P P P	Ca XVII Ca XIV Ca XV Fe XI Fe X Fe VIII
	186,600 186,850 186,880 187,890 188,230 190,040 191,050 192,040	P P P P P P	Fe VIII Fe XII Fe XII Fe XI Fe X Fe X	1	193.870 200.980 180.400 184.540 185.210	P P P P	Ca XIV Ca XV Fe XI Fe X Fe VIII
	186.850 186.880 187.890 188.230 190.040 191.050 192.040	P P P P P	Fe XII Fe XII Fe XII Fe XI Fe X		200:980 180.400 184.540 185.210 185.500	p p p p	Ca XV Fe XI Fe X Fe VIII
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	188,230 190,040 191,050 192,040	p P P P	Fe XXI Fe XI Fe X		184.540 185.210	p p	Fe X Fe VIII
	188.230 190.040 191.050 192.040	P P P	Fe XI Fe X		185.210	p	Fe VIII
	190.040 191.050 192.040	P P	Fe X		106 600		
	191.050 192.040	p	Po VIT		100,000	p	Fe VIII
	192.040		TO VIT		186.850	p	Fe XII
	100.000	p	Fe XXIV		186.880	p	Fe XII
	192.390	р	Fe XII		187.890	p	Fe XXI
	193.520	p	Fe XII		188.230	p	Fe XI
	194.660	p	Fe VIII		190.040	p	Fe X
	195.120	c	Fe XII		191.050	p	Fe XII
	195.970	p	Fe VIII		192.040	p	Fe XXIV
	196.540	p	Fe XIII		192.390	p	Fe XII
	196.650	p	Fe XII		193,520	p	Fe XII
	200.020	p	Fe XIII		194,660	p	Fe VIII
-	201.130	p	Fe XIII		195.120	c	Fe XII
	202.040	p	Fe XIII		195.970	p	Fe VIII
	203.830	p	Fe XIII		196.540	p	Fe XIII
	204,650	p	Fe XVII		196.650	p	Fe XII
	247,5160	p	Fe XXII		200.020	p	Fe XIII
	251.070	p	Fe XVI		201.130	p	Fe XIII
	251,960	p	Fe XIII		202.040	p	Fe XIII
	253,160	p	Fe XXII		203,830	p	Fe XIII
	254,870	p	Fe XVII		204.650	p	Fe XVII
	255.100	p	Fe XXIV		247.160	p	Fe XXII
	2	202.040 203.830 204.650 251.070 251.960 253.160 254.870 255.100 Sort: As	202.040 p 203.830 p 204.650 p 204.650 p 251.070 p 251.960 p 253.160 p 255.100 p 255.100 p Sort: Ascending	202.040 p Fe XIII 203.830 p Fe XIII 204.650 p Fe XVII 204.650 p Fe XVII 204.650 p Fe XVII 204.650 p Fe XVII 251.070 p Fe XVII 251.960 p Fe XVII 253.160 p Fe XVII 255.100 p Fe XVIV Sort: Ascending name Add	202.040 p Fe XIII 203.830 p Fe XIII 204.650 p Fe XVII 204.650 p Fe XVII 204.650 p Fe XVII 205.070 p Fe XVI 251.070 p Fe XVI 253.160 p Fe XXII 254.870 p Fe XXIV Sort: Ascending name Add	202.040 p Fe XIII 195.970 203.830 p Fe XIII 196.540 204.650 p Fe XVII 200.020 251.070 p Fe XVII 201.130 251.960 p Fe XVII 203.830 253.160 p Fe XVII 203.830 254.870 p Fe XVIV 204.650 255.100 p Fe XVIV 247.160 Sort: Ascending name Sort: Add	202.040 p Fe XIII 195.970 p 203.830 p Fe XIII 196.540 p 204.650 p Fe XVII 196.650 p 204.650 p Fe XVII 196.650 p 207.070 p Fe XVII 200.020 p 251.070 p Fe XVII 201.130 p 253.160 p Fe XVII 203.830 p 254.870 p Fe XVII 204.650 p 255.100 p Fe XXIV 247.160 p Sort Ascending name Sort Ascending

In all the cases above, you can add the lines you've chosen by hitting **Add**. **Dismiss** will get rid of the window without adding any lines (unless you've double-clicked on a line).

Adding a Custom Line

There will be certain cases where you want to add a line which isn't in the EIS project list, for example:

- the line you want isn't there;
- you want to include more than one line, or a piece of continuum, so the notation needs to be a custom one.

In any of these cases, the easiest thing to do is to add a project line and then customise the wavelength accordingly. (Remember that *the line list only contains the wavelengths*, not how wide the spectral windows are._)

In this example, we'll start by adding the s x project line at 180.73 Å, following the steps above.



We'll now move the central wavelength to 180.20 Å:



...and adjust the width of the window to 80 pixels, to create a wider spectral window than the four others:



And finally, we rename the window. This is done by going to the window's Name field (12) and clicking on Text (custom).



Here, you enter a name for the window, up to 8 characters long (this can include spaces). In this case, I've chosen "S X Wide", as I've just widened the window, but as long as the name is self-explanatory (you mightn't be the only one using this data, remember!), it should be fine.



Removing a Line from the Line List

If you want to remove a line, you have two options:

- from the main Linelist menu in (1), choose Linelist, Remove selected line from list; or
- right-click on the line in (7) and choose *Remove line* :

We can now select the custom line we created above, and right-click on it to bring up the *Remove* option. Note that this will only work on the selected (i.e. highlighted) line, not simply on whichever line you right-click on.



Saving a Line List

Once you've settled on a line list, you can save it for future use so that you don't have to keep entering the same parameters over and over every time you want to create a study. To save a line list, once it's been created, go to the main Linelist menu in (1), and choose Save in Database:



You'll then be presented with a dialogue box to fill in the information, which has some default information — you'll need to replace this, of course:

×	EIS Line List: Save	e to Database
Acronym:	[Core_Lines	(20 chars max / no spaces)
Author:	Matthew Whillock 😑	
Entry:	 Create new Update existing 	
Save to datab	ase Dismiss	

If your name is in the existing list of people associated with making EIS studies, then you can select it from the drop-down menu:



Otherwise, select *Unknown* from the list and enter your own name. Please keep this the same from one time to the next, as we would like the linelist database (and others) to be searchable by the author's name.

×	EIS Line List:	: Save to Database	×
Acronym:	Example_Lisť	(20 chars max / no spaces	0
Author:	Unknown	Arthur Guinness	
Entry:	 ◆ Create new ◇ Update existing 		
Save to data	pase Dismiss		

Once you've filled in this information, choose Save and you'll be asked if you're sure.



If you are sure, then click Yes. (The converse, naturally, also applies...)

Exporting a Line List

To submit a line list to the SSCs for inclusion in the official database, you need to export your choice of wavelengths to an ASCII text file. This is done by going to the main menu (1) again and choosing *Linelist, Export...* You'll then be presented with a dialogue box asking you for an **Acronym** for the line list, and an author (as for saving a linelist, above).

×	EIS Line List: Exp	ort	
Acronym:	Example_Lines	(20 chars max)	
Author:	David Williams 😑		
Export	Dismiss		

When you select *Export...*, you'll be asked where to save the linelist file (which, by default, is given a .def extension). The Raster Tool will come up with a filename based on your acronym, but you can change this to whatever you want.

Directory		
/home/drw/į		
Filter	Files	
*	sswid16,2,2207 sswid16,2,27414	A 1
Directories	sswid16.2.27886	
	SSW1016,2,28175 SSW1016,2,30494	
++ ObjSuite	sswid16.2.31395	
.Skype	sswid16.2.4337	
.TeXmacs	M 2	
Selection		
[/home/drw/examp]	le_lines_1.def	

Finally, click OK to save your list to the specified file.



Exit Linelit Rater Piol Edit Help Line List		EIS	Make Raster Tool		= 0 ×
Line List Verelength (Å) Name Pixel OCD Vixel Selected Line 10: 000000 1 192,80 0:0 Pervit 1199 1 22 10: 000000 1 200,90 Pervit 1199 1 22 10: 00000 1 200,90 Pervit 1199 1 22 10: 00000 1 200,90 Pervit 199 22 Pervit 1939 1955 1570 10: 00000 1 10: 10: Pervit 192 Pervit 1939 1955 201.32 Author: David David Pervit 10: 20 <td< th=""><th>Exit Linelist Raster Plot Edit</th><th></th><th></th><th></th><th>Help</th></td<>	Exit Linelist Raster Plot Edit				Help
Ouliet Active Flare Colour Table: RED-PURPLE 100 200 300 400 500 e0d 70 800 100 100 120 1540 1640 <th>Line List # ID: 000000 1 Acronym: Example_Lines Created: 20-Jun-2006 1 Author: David</th> <th>Wavelength(Å) Name 192.82 (c) Ca XVII 195.12 (c) Fe XII 200.98 (p) Ca XV 256.32 (c) He II</th> <th>Pixel CCD Width 1189 1 32 1292 1 32 1555 1 32 472 2 32</th> <th>Selected Line Pixels: 1539 : [1555 Wavelength: 200.63 : [200.98 Width: 32 </th> <th>: 1570 : 201.32 Full Range Ca XV</th>	Line List # ID: 000000 1 Acronym: Example_Lines Created: 20-Jun-2006 1 Author: David	Wavelength(Å) Name 192.82 (c) Ca XVII 195.12 (c) Fe XII 200.98 (p) Ca XV 256.32 (c) He II	Pixel CCD Width 1189 1 32 1292 1 32 1555 1 32 472 2 32	Selected Line Pixels: 1539 : [1555 Wavelength: 200.63 : [200.98 Width: 32	: 1570 : 201.32 Full Range Ca XV
100 200 300 400 500 600 700 800 900 1000 1100 1200 1500 1600 1700 1800 1900 2000 Slit/slot selection 1 Raster Type: Scanning Slit'n'Stare Window height (pixels) 512 Default width (pixels) 32 Set all Number of exposures/step 1 Set Exposure/Delay Times Exposure(Delay) time s(ms) 50.00(10) Number of Fine Mirror steps 140 Step size 2 (Number of exposures = 41) Raster ID: 000001 Acronym: Default Raster Author: John Rainnie Created: 26-May-2006 Duration: 35m33s Volume: 41984 kBits Data rate: 20.1484 kBits/s Tite: (RA:000001] [Scan:40 (2*) steps] [s:1*] [wH:512, nWins:4, LL:000000] [ExpT(Delay):50.00 (10)] Solar-B EISS	Quiet Active Flare Colour Table 100 200 300 400 500 170 172 174 176 13	e: RED-PURPLE 606 70 800 90 78 190 182 184 186	0 1000 1100 1200 188 190 192 194	1300 1400 1500 1400 1700 196 198 200 202 204	1800 1900 2000 206 208 210
Number of exposures/step 1 Set Exposure/Delay Times Exposure(Delay) time s(ms) 50.00(10) Number of Fine Mirror steps 40 Step size 2 (Number of exposures = 41) Raster ID: 000001 Acronym: Default Raster Author: John Rainnie Created: 26-May-2006 Duration: 35m33s Volume: 41984 kBits Data rate: 20.1484 kBits/s EIS EIS Title: [RA:000001] [Scan:40(2*) steps][ss:1*] [wH:512, nWins:4, LL:000000] [ExpT(Delay) :50.00(10)] Created: Solar-B EIS	100 200 300 400 500 Slit/slot selection 1	600 700 800 90 Raster Type: Scanning Default width (pixe	a 1000 1100 1200 Sit'n'Stare	1300 1400 1500 1600 1700 Set all	1800 1900 2000
Raster ID: 000001 Acronym: Default Raster Author: John Rainnie Created: 26-May-2006 Duration: 35m33s Volume: 41984 kBits Data rate: 20.1484 kBits/s EIS Solar-B Title: [RA:00001][Scan:40(2*)steps][ss:1*][wH:512, nWins:4, LL:000000][ExpT(Delay):50.00(10)] 50.00(10)] 50.00(10)] 50.00(10)]	Number of exposures/step 1 S Number of Fine Mirror steps 40 Step	et Exposure/Delay Times Exp p size 2 (Number of expo	osure(Delay) time s(ms) 50.00 osures = 41))(1 0)	
	Raster ID: 000001 Acronym: D Duration: 35m33s Volume: 4 Title: [RA:000001] [Scan:40(2*)]	efault Raster Author 1984 kBits Data ra)steps][ss:1"][wH:512,nWins	: John Rainnie (ate: 20.1484 kBita/s :4.LL:000000][ExpT(Delay)	Created: 26-May-2006 :50.00(10)]	EIS
Exit			Exit		

Importing a Line List

This is just the inverse of exporting a linelist, and is useful if you want to see a linelist that someone else has created, but that isn't necessarily in the official database. It's pretty straight-forward, but I'll fill in the details on this later.

Changing the CCD image

Before launch, only synthetic CCD (i.e. full-spectrum) images are available, and this is what this guide will feature until the post-commissioning phase.

(Insert a bit more chat about this)

- Active Region & Flare spectra
- SERTS already-observed spectra?
- EIS commissioning spectra (spectral atlas)

Note that when you change colour tables, the spectrum silently defaults back to the Quiet Sun image -- you need to change this manually to make sure you're looking at the right kind of spectrum.

The spectra which are currently available are CHIANTI-derived synthetic spectra, and there are three types from which to choose:

- 1. Quiet (Quiet Sun)
- 2. Active (Active Region)
- 3. Flare

For now, try changing the type from Quiet to Active.



You can also change the colour table in order to see fainter (or brighter) lines more clearly. For example, the Hardcandy table:



Making a Raster

Changing the type of raster

This is easy enough to do, but it has a knock-on effect. For the purposes of this example we'll work with a Sit'n'Stare raster. To do this, go to the *Raster Type* section (19) and choose *Sit'n'Stare*. You'll notice that some of the fields below change, namely:

Scanning Raster Sit'n'Stare Raster

Number of exposures/step Number of exposures/set

Number of Fine Mirror steps Number of sets

Change the type of raster to Sit'n'Stare:

	EIS Make Raster Tool	
xit Linelist Raster Plot Edit		He
Line List ID: 000000 Acronym: Example_Lines Created: 20-Jun-2006 1 Author: David	# Wavelength (Å) Name Pixel OCD Width Selected Line 1 192.82 (c) Ca XVII 1189 1 32 2 195.12 (c) Fe XII 1292 1 32 4 205.32 (c) He II 472 2 32 Width: 32	5 : 1570 98 : 201.32 Full Range
 Quiet Active Flare Colour Ta 100 200 300 400 5 170 172 174 176 100 200 300 400 5 	Ible: RED-PURPLE 00 600 70 800 900 1000 1100 1200 1300 1400 1560 1600 178 180 182 184 186 180 192 194 196 198 200 202 00 600 700 800 900 1000 1100 1200 1300 1400 1500 1600	1700 1800 1900 2000 204 206 208 210
it/slot selection 1 = indow height (pixels) 512 umber of exposures/set 1 = umber of exposure sets 1	Raster Type: Scanning Sit'n'Stare Default width (pixels) 32 Set all Set Exposure/Delay Times Exposure(Delay) time s(ms) 50.00(10) Duration (s) [50.010]	
Raster ID: 000001 Acronym: Duration: 53s Volume: Title: [RA:000001] [SNS:1se	Default Raster Author: John Rainnie Created: 26-May-2006 1024 kBits Data rate: 19.6569 kBits/s t][ss:1*][vH:512,nWins:4,LL:000000][ExpT(Delay):50.00(10)]	EIS
	Evit	

This is simply logical, and we'll explain the terminology as we go on.

Making a Study

Notes

Just to keep track of things, I've appended some notes below.

- Only one response study possible per trigger type, between real-time (i.e. commanding) contacts
 - *E.g.*, If two studies in the timeline are sensitive to the EIS flare trigger, they must both respond with the same response study. There is only room for one set of EIS Flare Trigger response properties in the EIS ICU.

The same goes for the XRT flare trigger, EIS event trigger and the Automatic Exposure Control.

-- DaveWilliams - 20 Jun 2006

* Change the height of the slit image:



* Change the width of all the spectral windows:

	EIS Make Raster T	ool	
xit Linelist Raster Plot Edit			Hel
Line List # ID: 000000 1 Acronym: Example_Lines Created: 20-Jun-2006 1 Author: David	Wavelength(Å) Name Pixel CCD 192.82 (c) Ca XVII 1189 1 195.12 (c) Fe XII 1292 1 200.98 (p) Ca XV 1555 1 256.32 (c) He II 472 2	Width Selected Line 72 Pixels: 1519 : [1] 72 Wavelength: 200.18 : [2] Width: 72 Wavelength: 200.18 : [2] Width: 72	555 : 1590 00.98 : 201.76 Full Range otation Ca XV
 ◆ Quiet ◇ Active ◇ Flare Colour Tab 100 200 300 400 500 170 172 174 176 1 	e: RED-PURPLE	1100 1200 1300 1400 1500 1600	0 1700 1800 1900 2000 204 206 208 210
lit/slot selection 1	Raster Type: Scanning Sit'n'Stare Default width (pixels) 72 Set Exposure/Delay Times Exposure(Delay) time ration (s) [50.010	Set all ne s(ms) 50.00(10)	
Raster ID: 000001 Acronym: 1 Duration: 53s Volume: Title: [RA:000001] [SNS:1set])efault Raster Author: John Ra 1620 kBits Data rate: 31.1912 [ss:1*] [wH: 360, nWins:4, LL:000000] [ExpT (De	innie Created: 26-May-2006 kBits/s slay):50.00(10)]	EIS
	Exit		

* Changing the width of only the Ca XV window:

	EIS Make F	Raster Tool	
xit Linelist Raster Plot Edit			Hel
Line List ID: 000000 Acronym: Example_Lines Created: 20-Jun-2006 1 Author: David	# Wavelength(Å) Name Pixe 1 192.82 (c) Ca XVII 1189 2 195.12 (c) Fe XII 1292 3 200.98 (p) Ca XV 1555 4 256.32 (c) He II 472	1 CCD Width Se 9 1 72 2 1 72 5 1 95 2 72 W W Na	lected Line xels: 1507 : 1555 : 1602 avelength: 199.91 : 200.98 : 202.03 idth: 96
 Quiet Active Flare Colour Ta 100 200 300 400 5 170 172 174 176 100 200 300 400 5 	ble: RED-PURPLE	000 1100 1200 130 190 192 194 1 000 1100 1200 130	e 1400 1500 1600 1700 1800 1900 2000 196 198 200 202 204 206 208 210
Iit/slot selection 1 /indow height (pixels) 360 lumber of exposures/set 1 lumber of exposure sets 1	Raster Type: Scanning Si Default width (pixels) Set Exposure/Delay Times Exposure(D uration (s) \$0.010	It'n'Stare 72 Delay) time s(ms) 50.00(10)	Set all
Raster ID: 000001 Acronym: Duration: 53s Volume: Title: [RA:000001] (SNS:1set)	Default Raster Author. J 1755 kBits Data rate: 3][ss:1*][wH:360, nWins:4, LL:000000][John Rainnie Create 33.7905 kBita/s ExpT(Delay):50.00(10)]	d: 26-May-2006
		Exit	

* Changing the number of different exposure durations per exposure "set":



* Setting the different exposure times and their delays:

K	EIS Raster: S	et Exposi	ıre Times	
	Set Expo	sure T:	imes (s)	
Exposure 1:	Time (s):) 50.00	Delay (ms):	<u>]</u> 10
Exposure 2:	Time (s):	1į̇́0.00	Delay (ms):	1į̇́0
Exposure 3:	Time (s):	5 <u>,</u> 00	Delay (ms):	10
		Accept		

* The new exposure times are now set:

N	EIS Make Ra	ster Tool		
Exit Linelist Raster Plot Edit				Help
Line List # ID: 000000 1 Acronym: Example_Lines 2 Created: 20-Jun-2006 1 Author: David	Wavelength(Å) Name Pixel 192.82 (c) Ca XVII 1189 195.12 (c) Fe XII 1292 200.98 (p) Ga XV 1555 256.32 (c) He II 472	CCD Width Sa 1 72 F 1 72 F 2 72 V	elected Line Pixels: 1507 : [1555 : 16 Wavelength: 199.91 : [200.96 : 20 Width: 96 Warme: Spectral notation Ca Text (custom) Ca	i02 12.03 Full Range
 Quiet < Active Flare Colour Table 100 200 300 400 500 170 172 174 176 17 100 200 300 400 500 	RED-PURPLE 00 100 600 700 800 900 100 8 180 182 184 186 188 600 700 800 900 100	0 1100 -1200 -13 190 192 194	00 1400 15 <u>80 18</u> 00 1700 18 196 198 200 212 204 206	900 1900 2000 208 210 800 1900 2000
Ill/slot selection 1 /indow height (pixels) 360 lumber of exposures/set 3 lumber of exposure sets 1	Raster Type: Scanning Sit'n Default width (pixels) et Exposure/Delay Times Exposure(Del ation (s) [65.030	"Stare 72 ay) time s(ms) 50.00(10)	Set all ,10.00(10),5.00(10)	
Raster ID: 000001 Acronym: D Duration: 1m10s Volume: 5 Title: [RA:000001][SNS:1set][efault Raster Author: Job 265 kBits Data rate: 76. ss:1*][wH:360, nWins:4, LL:000000][Ex	nn Rainnie Creat 6880 kBits/s pT(Delay):50.00(10)10	ed: 26-May-2006	IS
		Exit		

* Change the number of sit'n'stare sets to 20:



* Saving the raster to the database:



* If the corresponding linelist hasn't yet been saved to the database ...:



* Exporting a raster to ASCII format:

K	EIS Line List: Exp	ort	
Acronym:	Example_Rasteř	(20 chars max)	
Author:	David Williams 😑		
Export	Dismiss		

* Specifying the raster's exported filename:

X	Please Select a File for Writing	
Directory /home/drw/į́		
Filter	Files	
*j Directories .AbiSuite .Skype .TeXmacs	sswid15,5,30919 sswid15,5,31522 sswid15,5,7699 sswid15,6,17269 sswid16,2,11296 sswid16,2,2207 sswid16,2,27414 sswid16,2,27886	
Selection Vhome/drw/e	xample raster 1.def	
	Filter	Cancel

* The finished product?:



This topic: SolarB > EISPlanningToolGuide

History: r22 - 20 Jun 2006 - 17:06:09 - <u>DaveWilliams</u>

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